

PIC simulation services at the CCMC in support of MMS mission

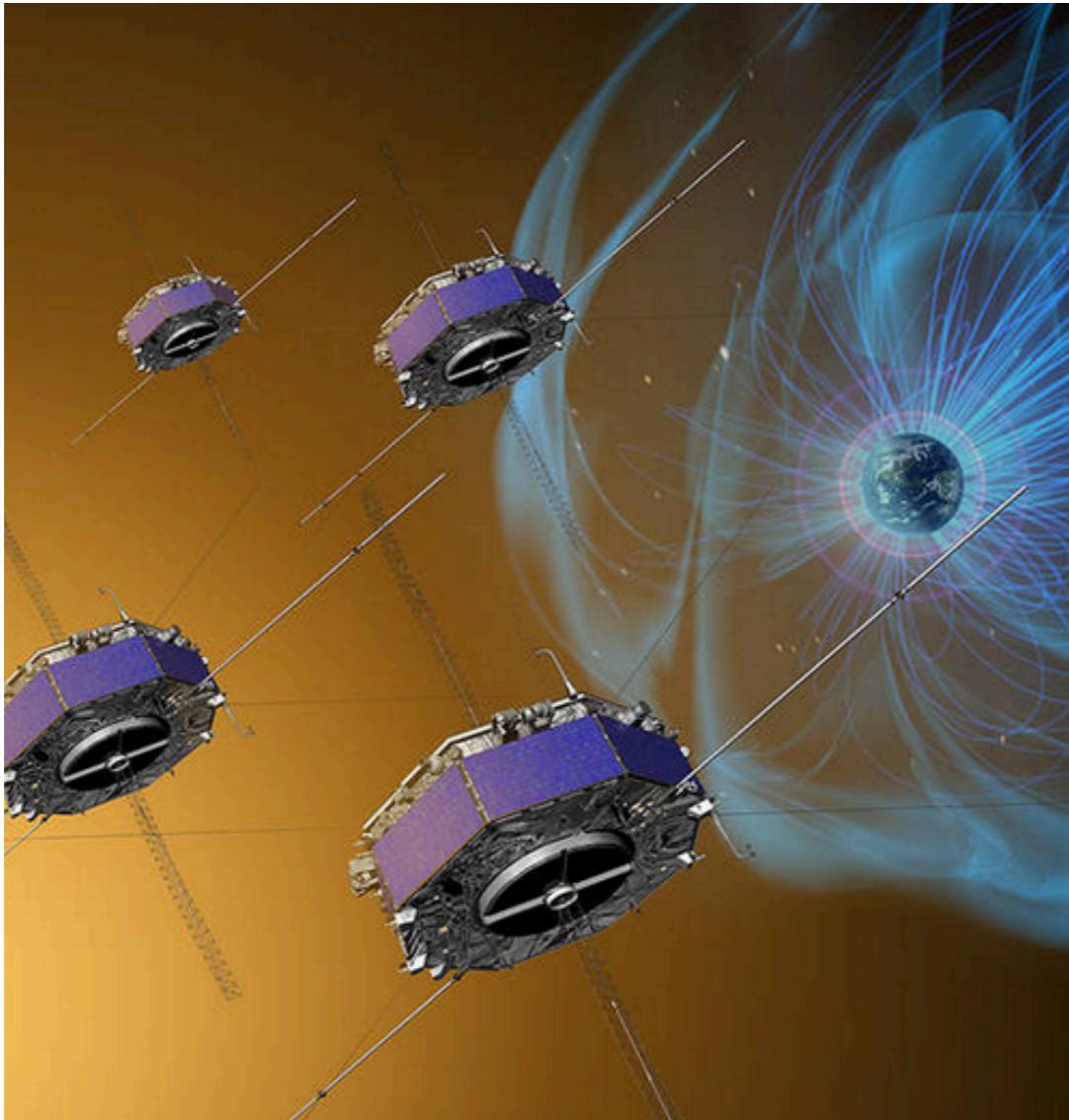
Yi-Hsin Liu
Masha Kuznetsova
Lutz Rastaetter
Asher Pembroke
Michael Hesse



The era of MMS

Magnetospheric Multiscale Mission (MMS)

March 13, 2015

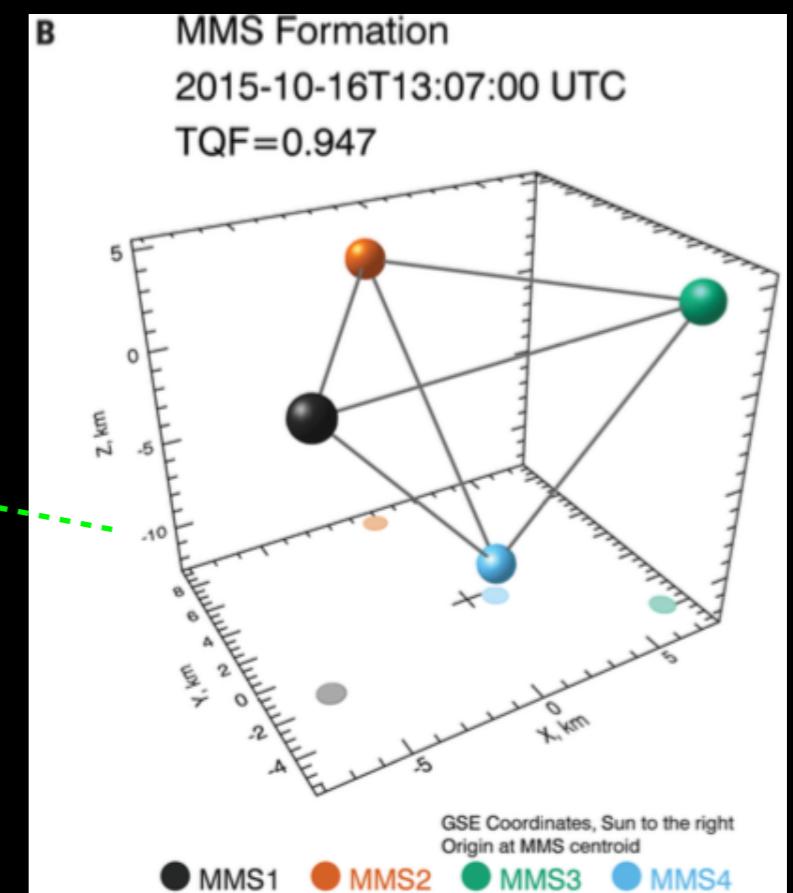
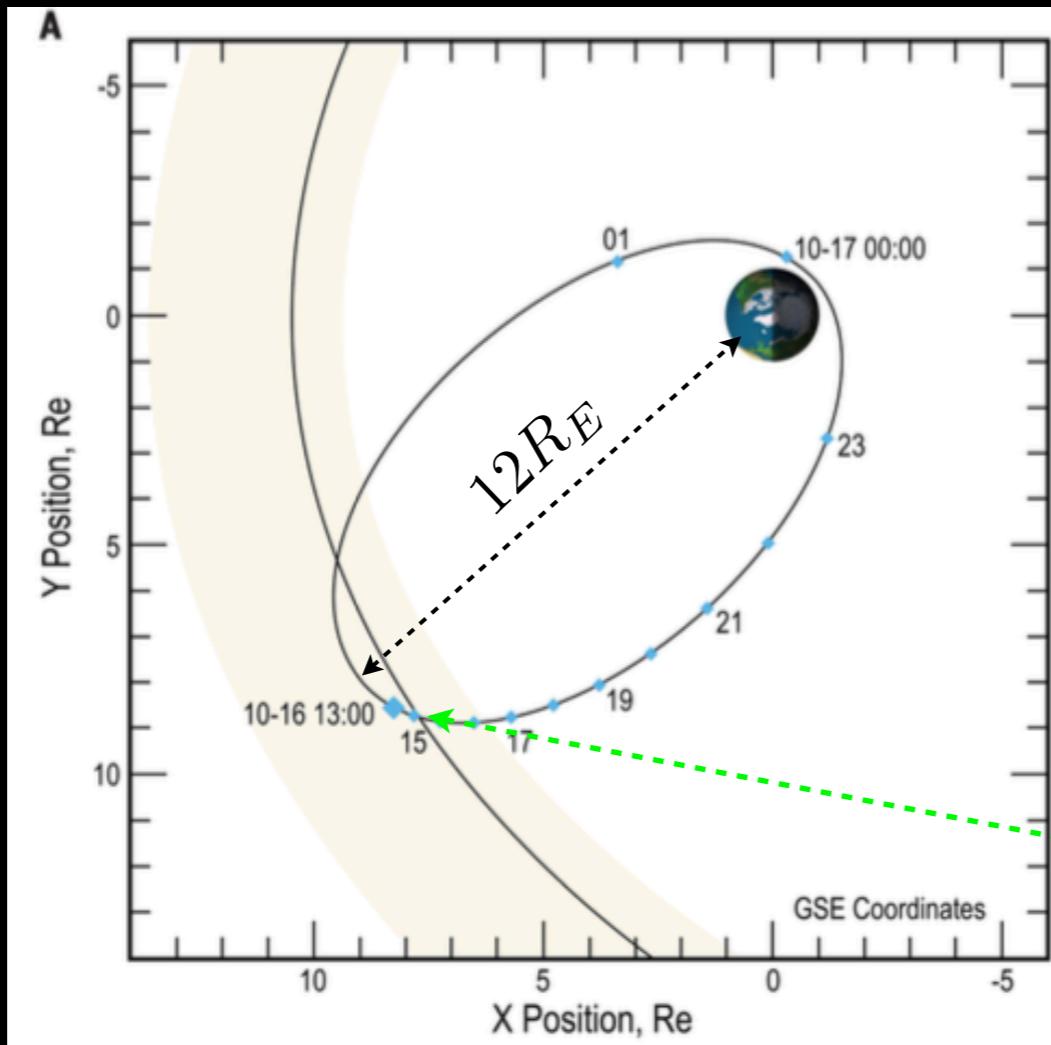
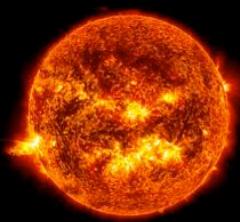


<http://mms.gsfc.nasa.gov>

tight tetrahedron formation: separation down to 7 km!
100x faster for electrons (30 ms)
30x faster for ions (150 ms)

- MMS leads us into a stage where the **electron-scale** physics of magnetic reconnection, in nature, can be resolved in an unprecedented manner!!
- PIC simulations can self-consistently model these electron-scale physics.

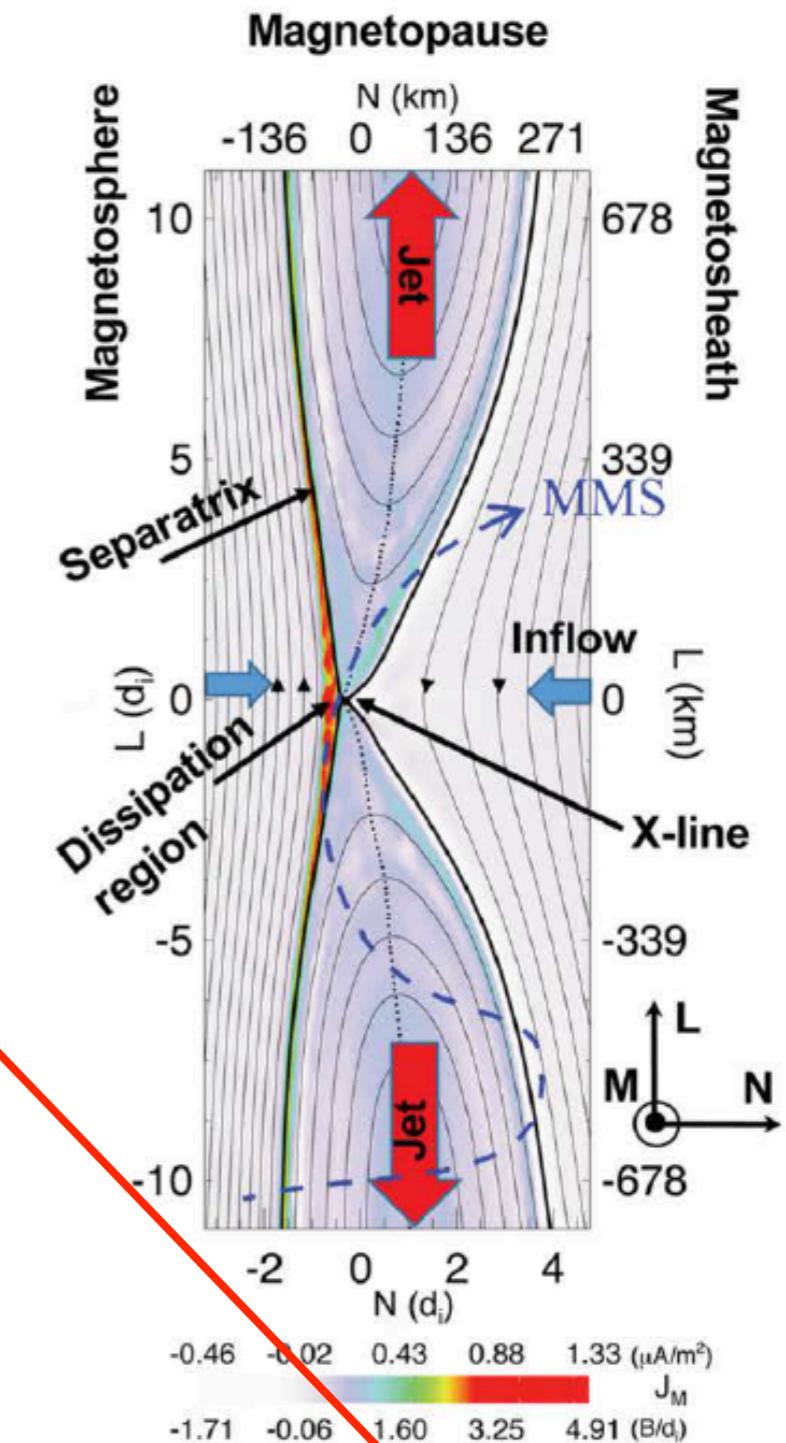
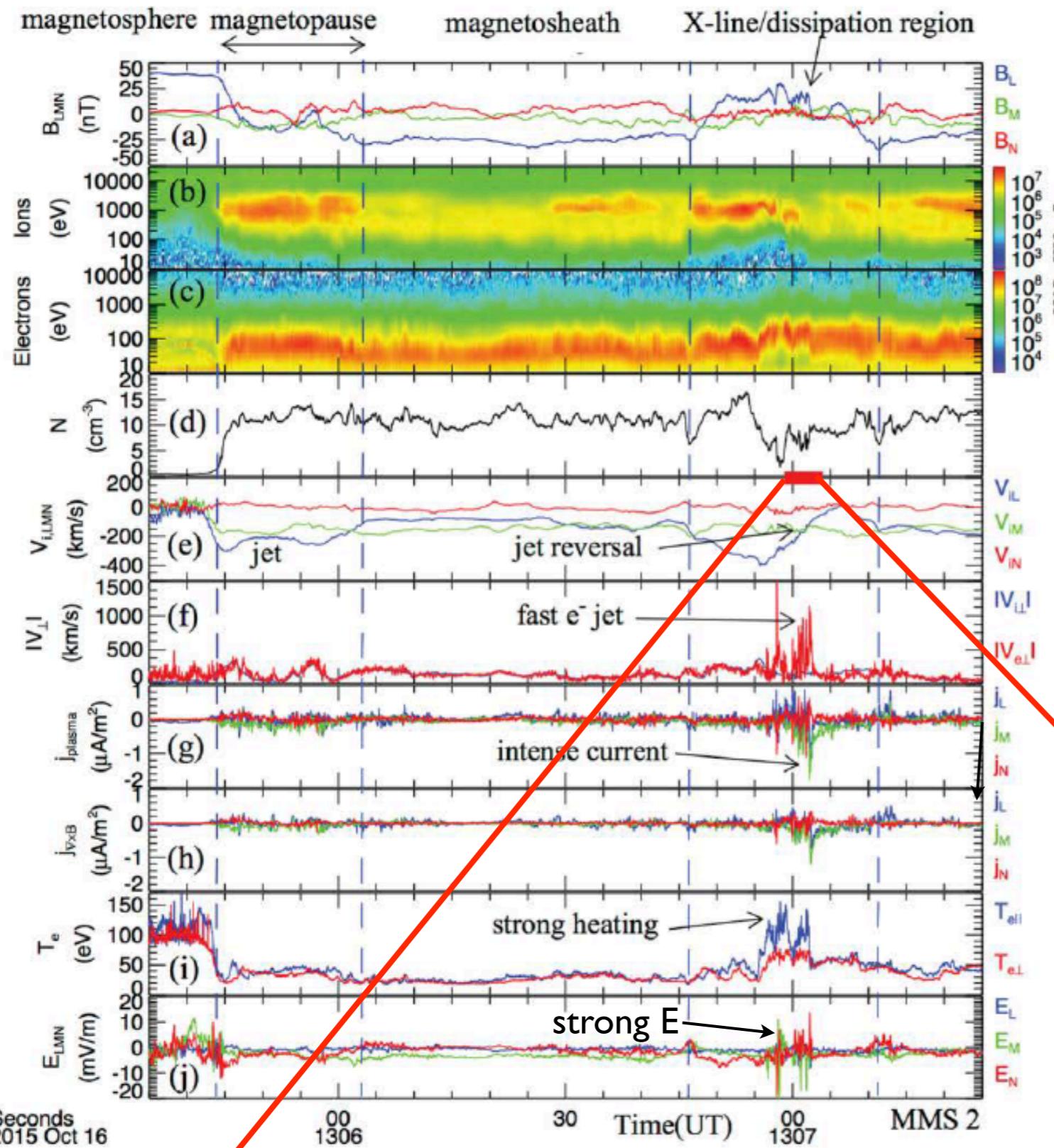
Spacecraft orbit on October 16, 2015



Encounter of the electron diffusion region!

MMS2

(Burch et al. SCIENCE 2016)



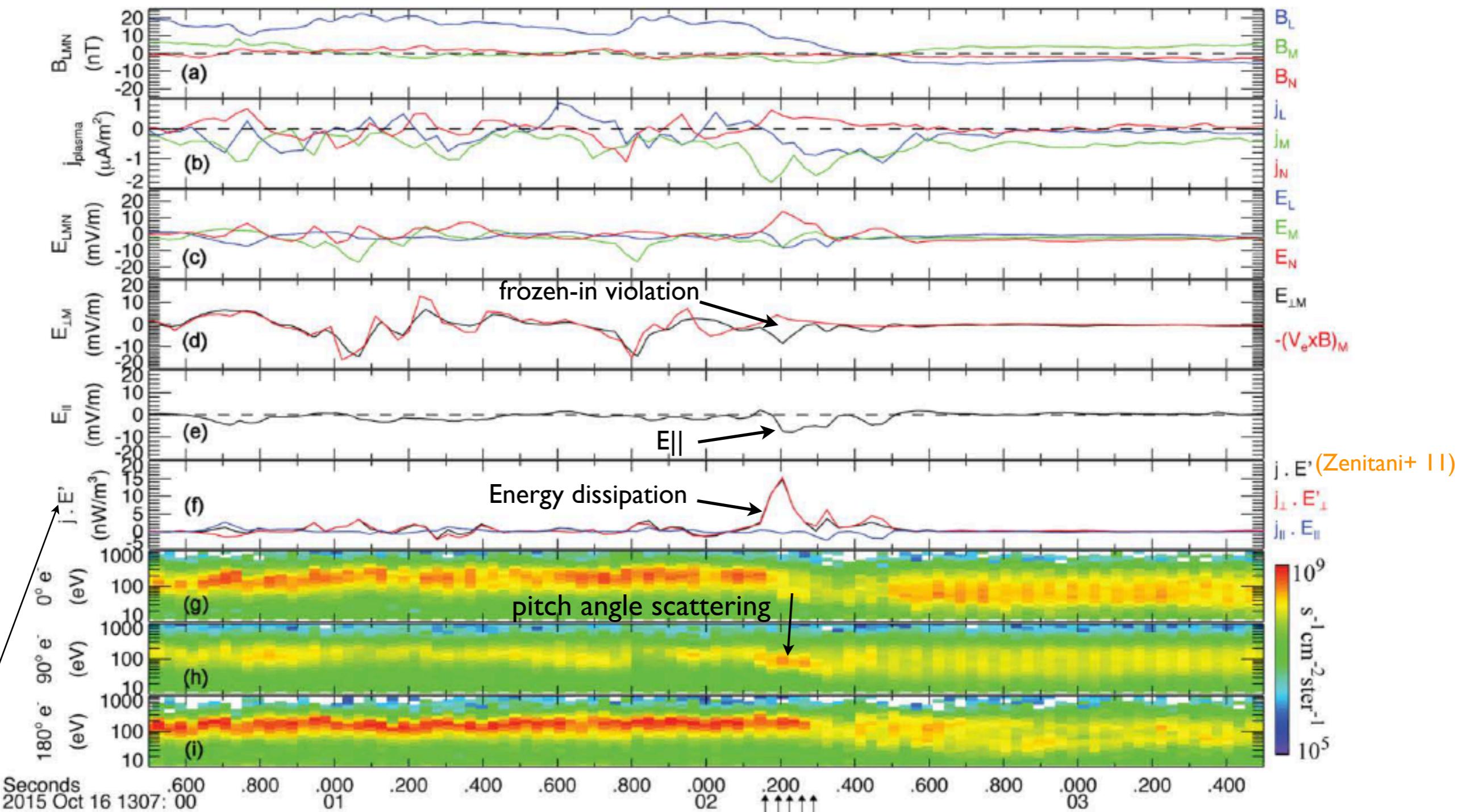
ER~ 1-5 mA/m

VA~ 400 km/s

vthe,sph~ 6,000 km/s; vthi,sph~ 300 km/s

More signatures of electron diffusion region...

MMS2



$$E' = E + V_e \times B$$

$ER \sim 1-5 \text{ mA/m}$

$\lambda_D \sim 60 \text{ m}$

$de \sim 1.66 \text{ km}; di \sim 70 \text{ km}$

$wce,sph \sim 1000 \text{ Hz}; wci,sph \sim 0.5 \text{ Hz}$

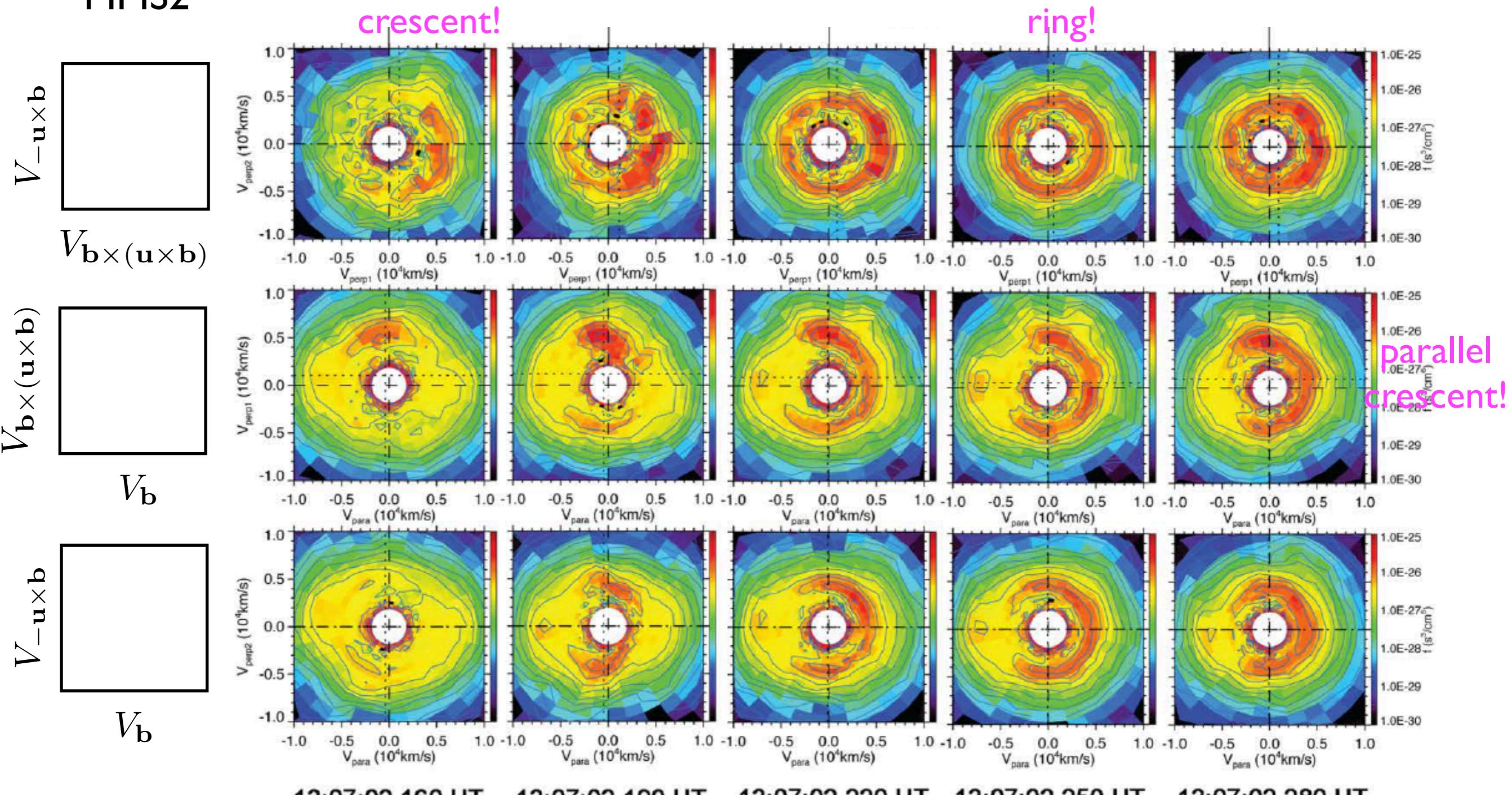
$wpe,sph \sim 10,000 \text{ Hz}; wpi,sph \sim 200 \text{ Hz}$

$VA \sim 400 \text{ km/s}$

$vthe,sph \sim 6,000 \text{ km/s}; vthi,sph \sim 300 \text{ km/s}$

Electron distribution inside reconnection diffusion region

MMS2

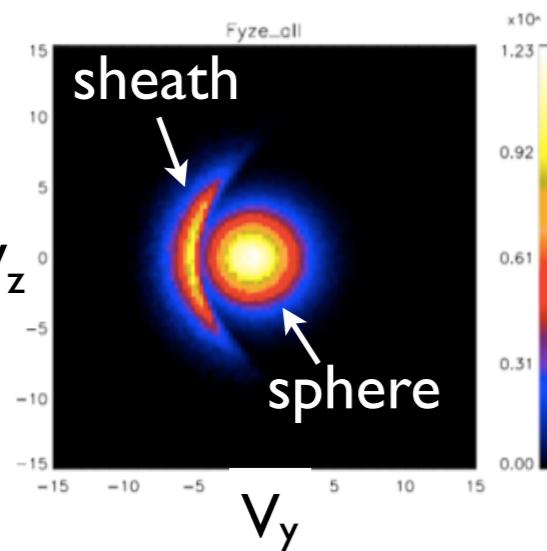


13:07:02.160 UT 13:07:02.190 UT 13:07:02.220 UT 13:07:02.250 UT 13:07:02.280 UT

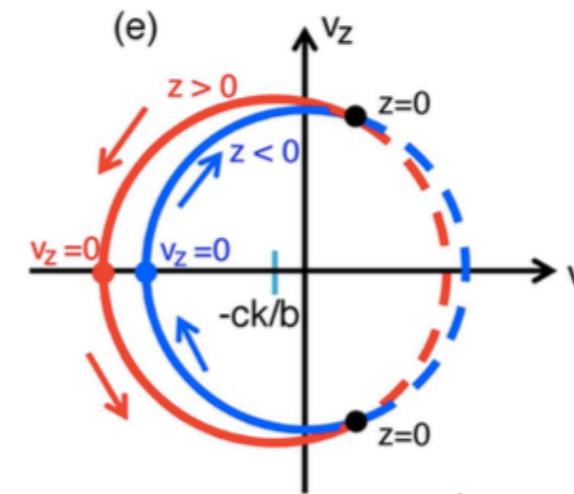
2015 October 16

One of the popular distributions: Crescent

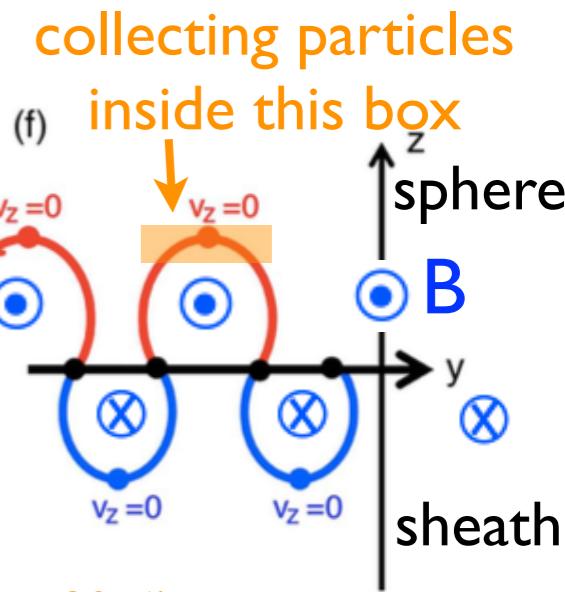
Reduced



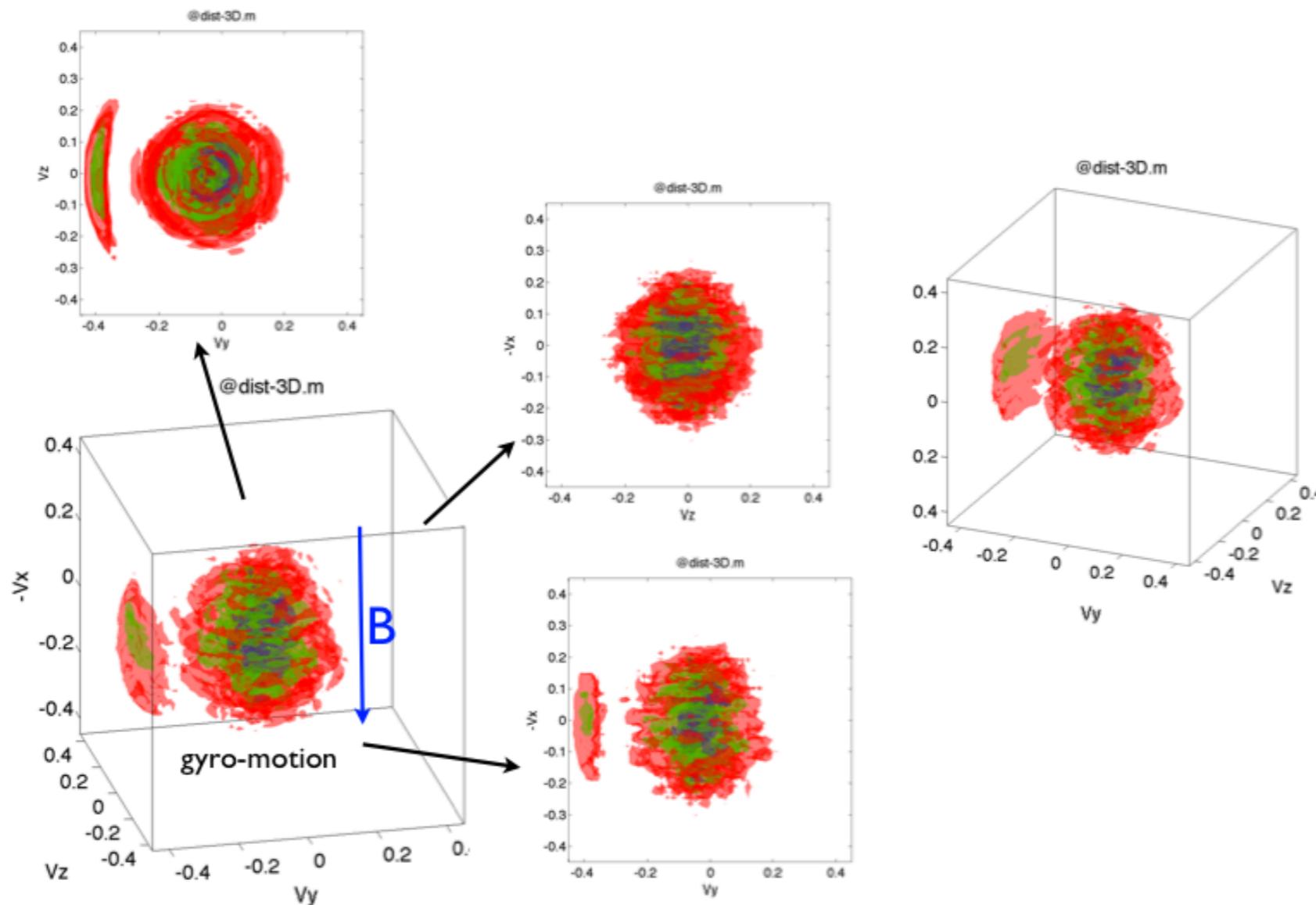
(Hesse et al. 2013;
Hesse et al. 2016;
Bessho et al. 2016;
Chen et al. 2016;
Shay et al. 2016;
Egedal et al. 2016)



(Bessho et al. 2016)

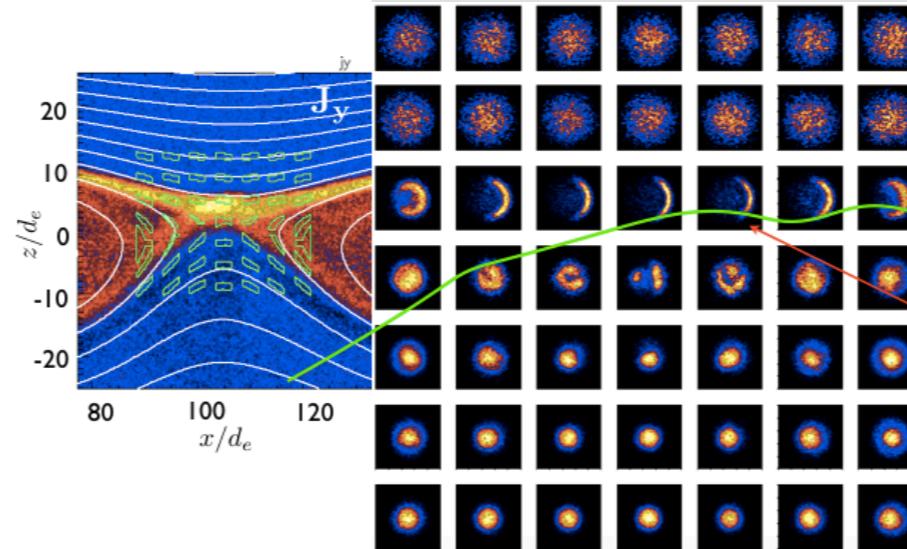


Full 3D



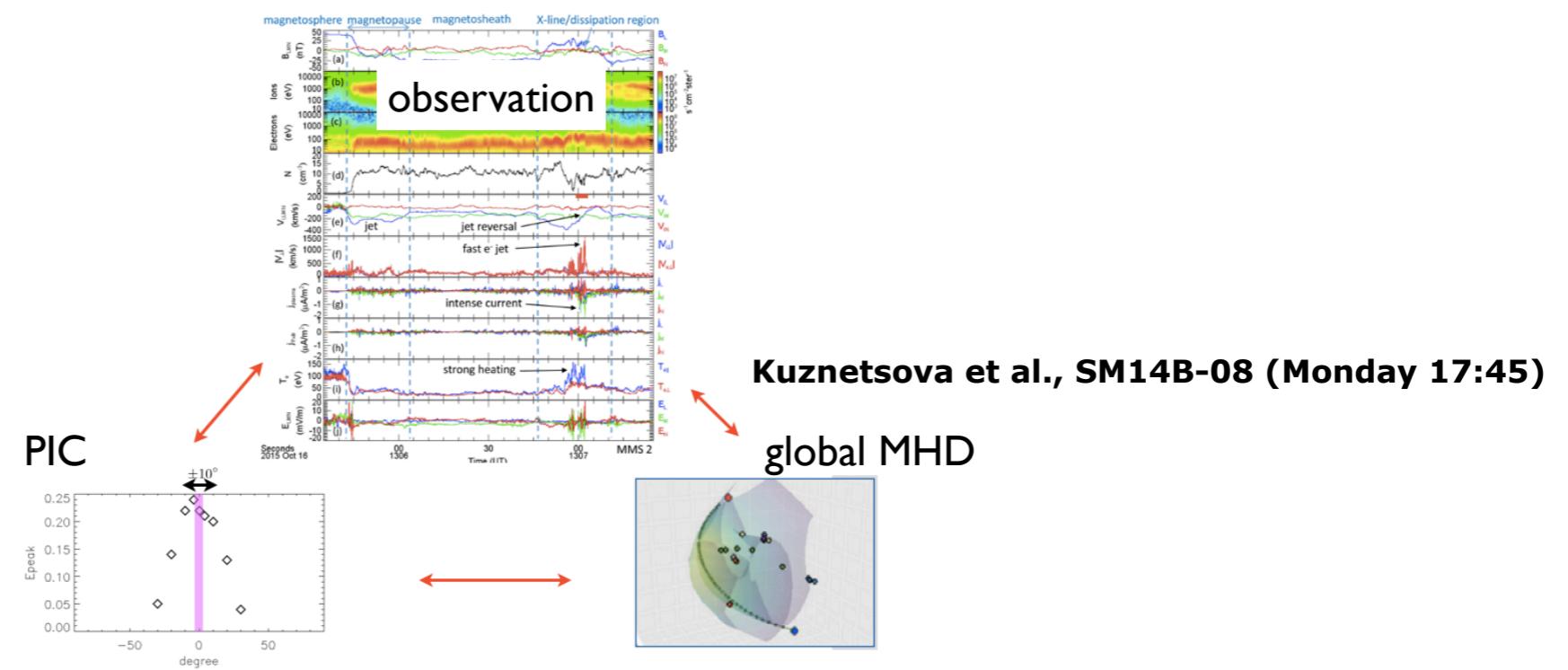
PIC services at the CCMC

I. Help interpret particle distributions at geometrically important locations.



2. Help determine the LMN coordinate.

- A potential science project that joints effort of MMS observation, local PIC & Global MHD simulations to evaluate the importance of local versus global physics in controlling the x-line orientation.

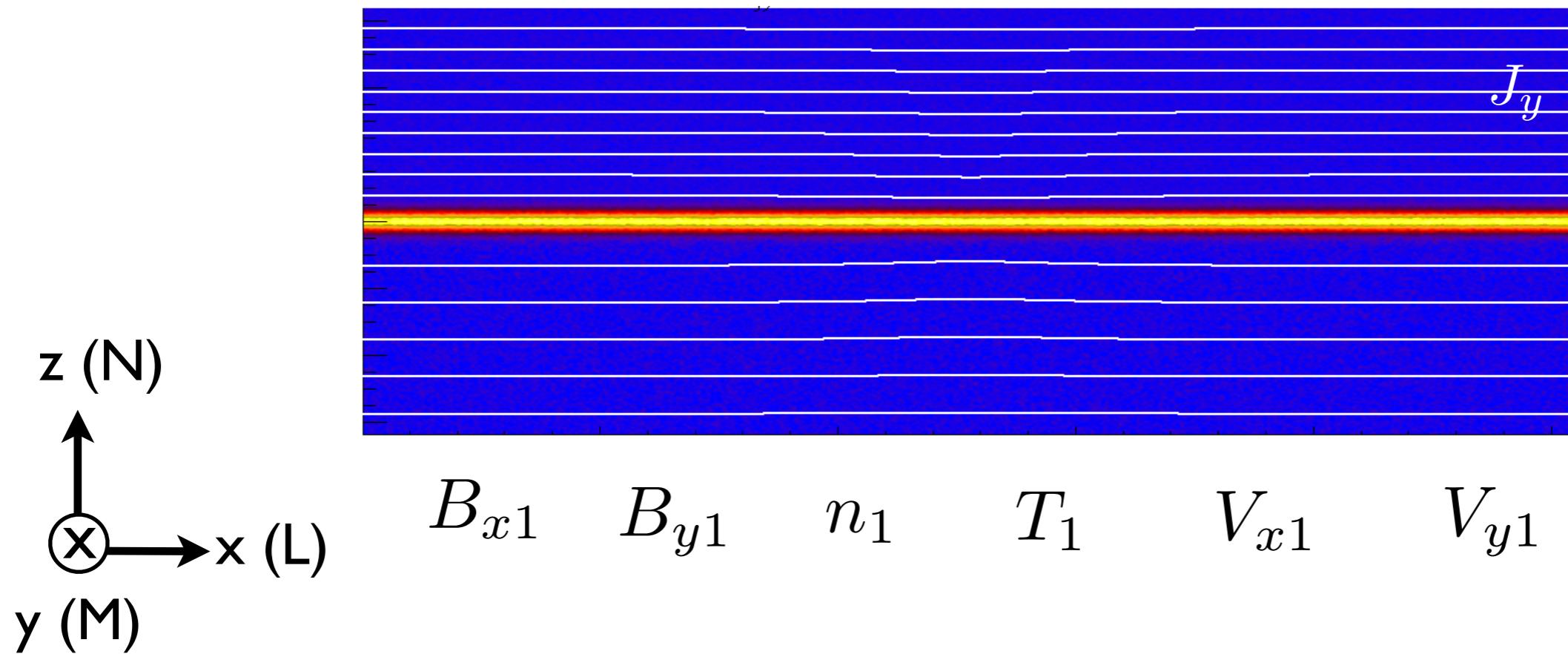


Service I:

Help interpret particle distributions at geometrically important locations.

We have generalized the initial condition

$$B_{x2} \quad B_{y2} \quad n_2 \quad T_2 \quad V_{x2} \quad V_{y2}$$

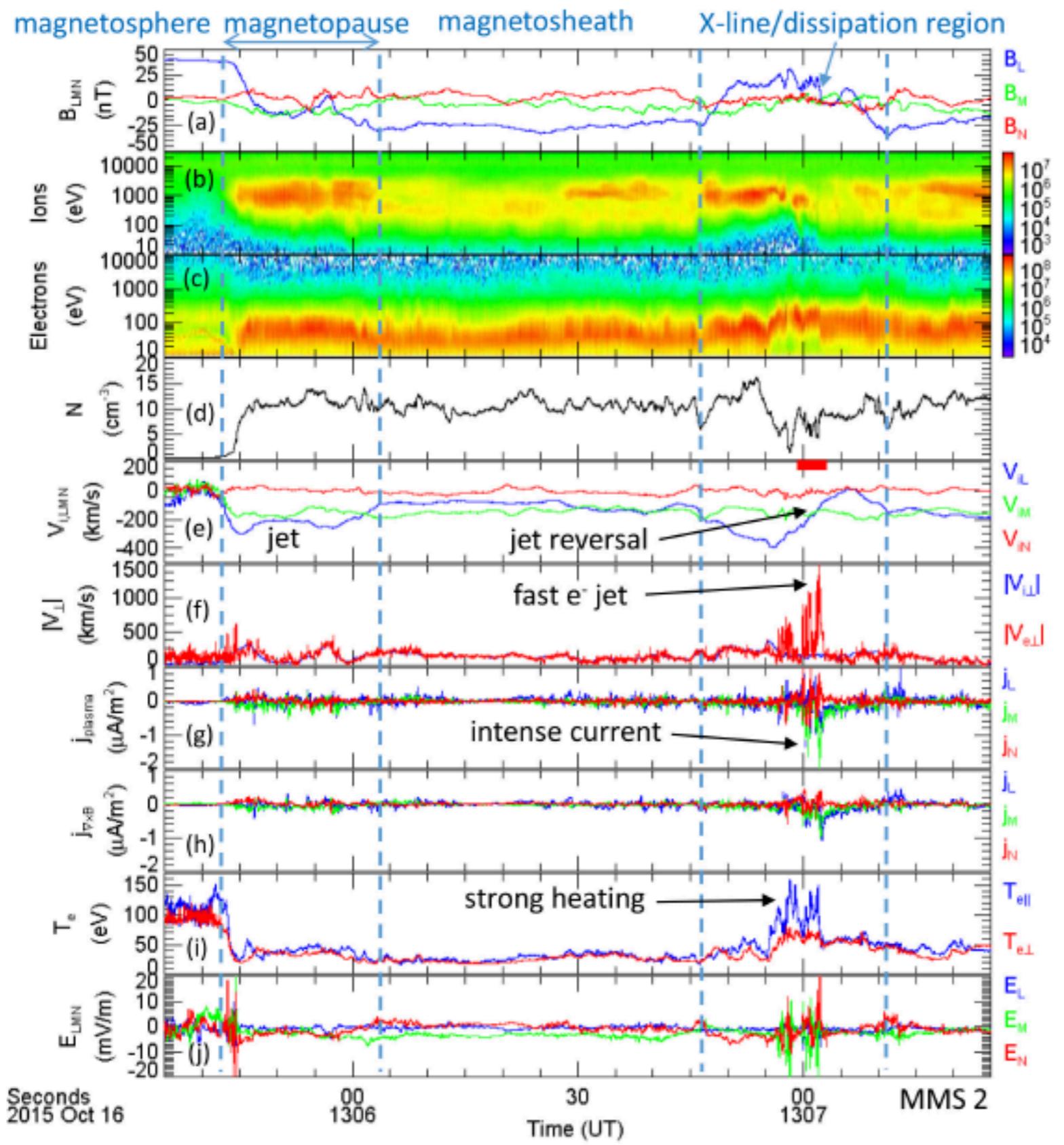


- The upstream conditions cover a wide range of variety.

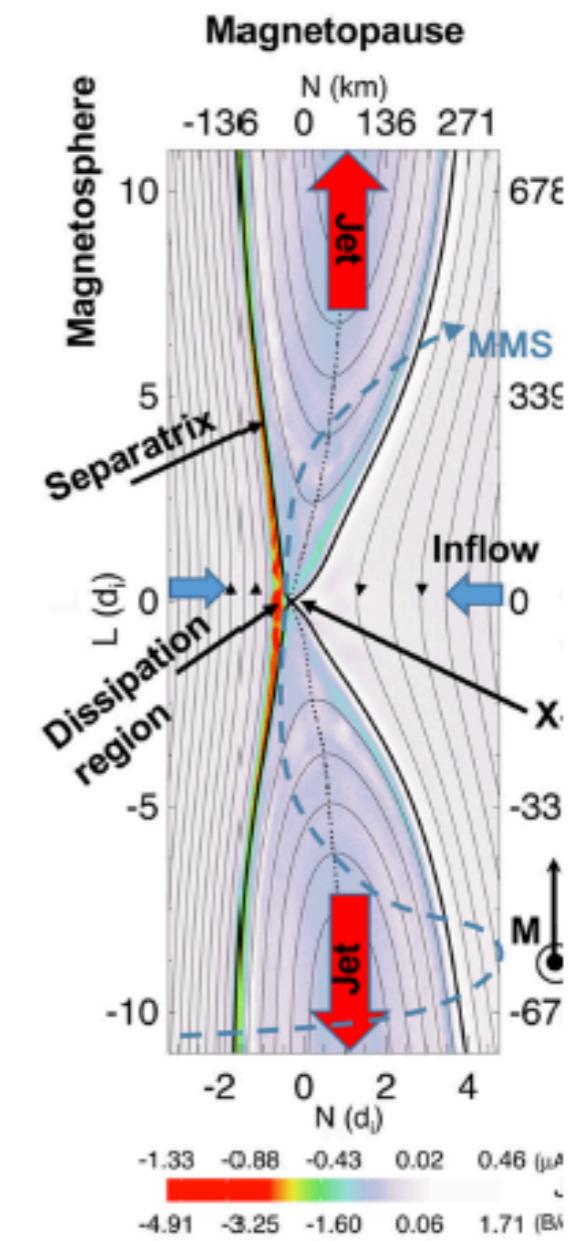
Step I: Give us the upstream condition

2

I



PIC simulation

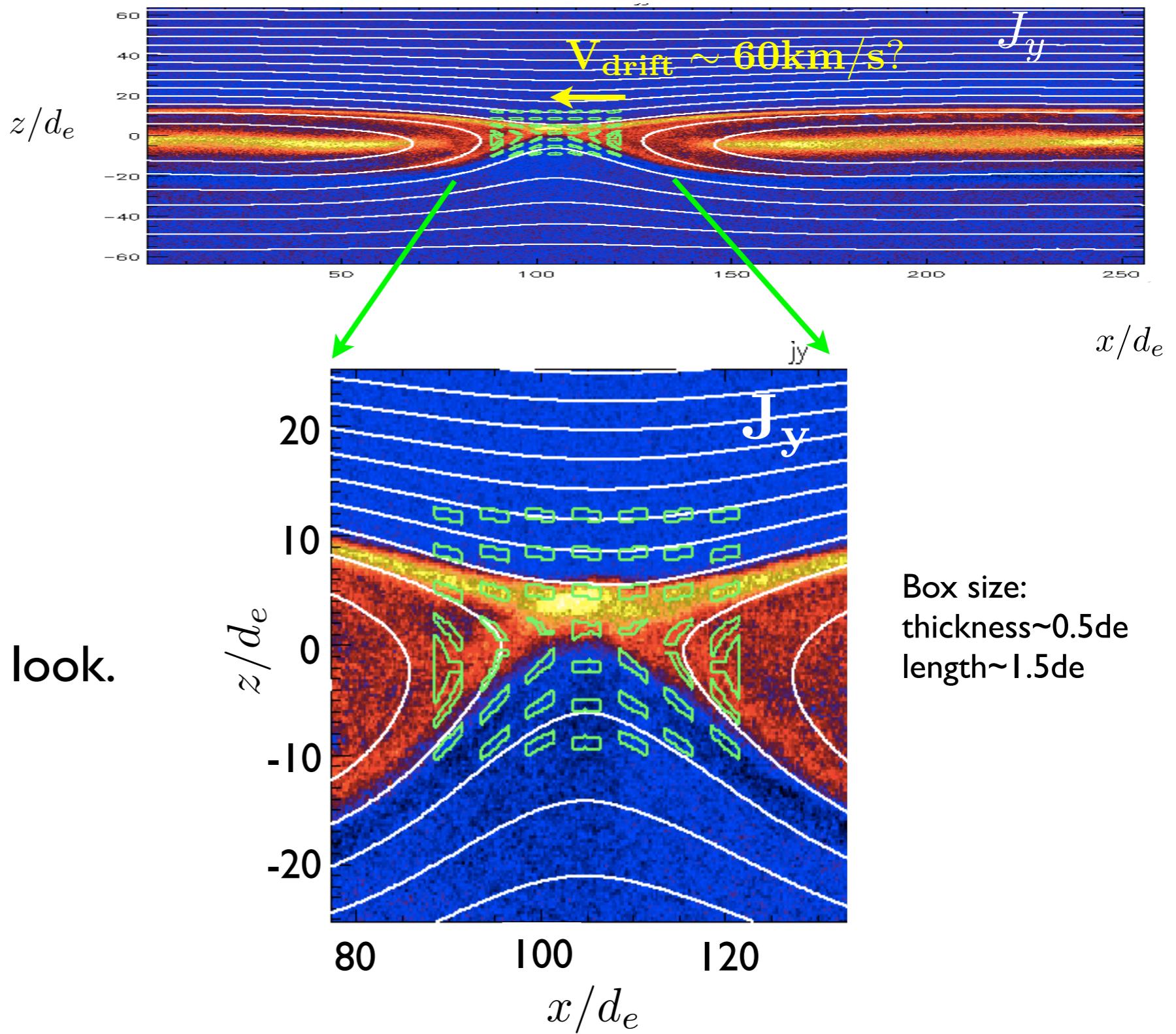


(Burch et al. SCIENCE 2016)

Step 2: We make a run, generate fields, moments & distributions

Quick-look:

- Field-aligned boxes.
- @ Geometrically important locations.
- $\sim 7 \times 7$ boxes should be enough for a quick look.



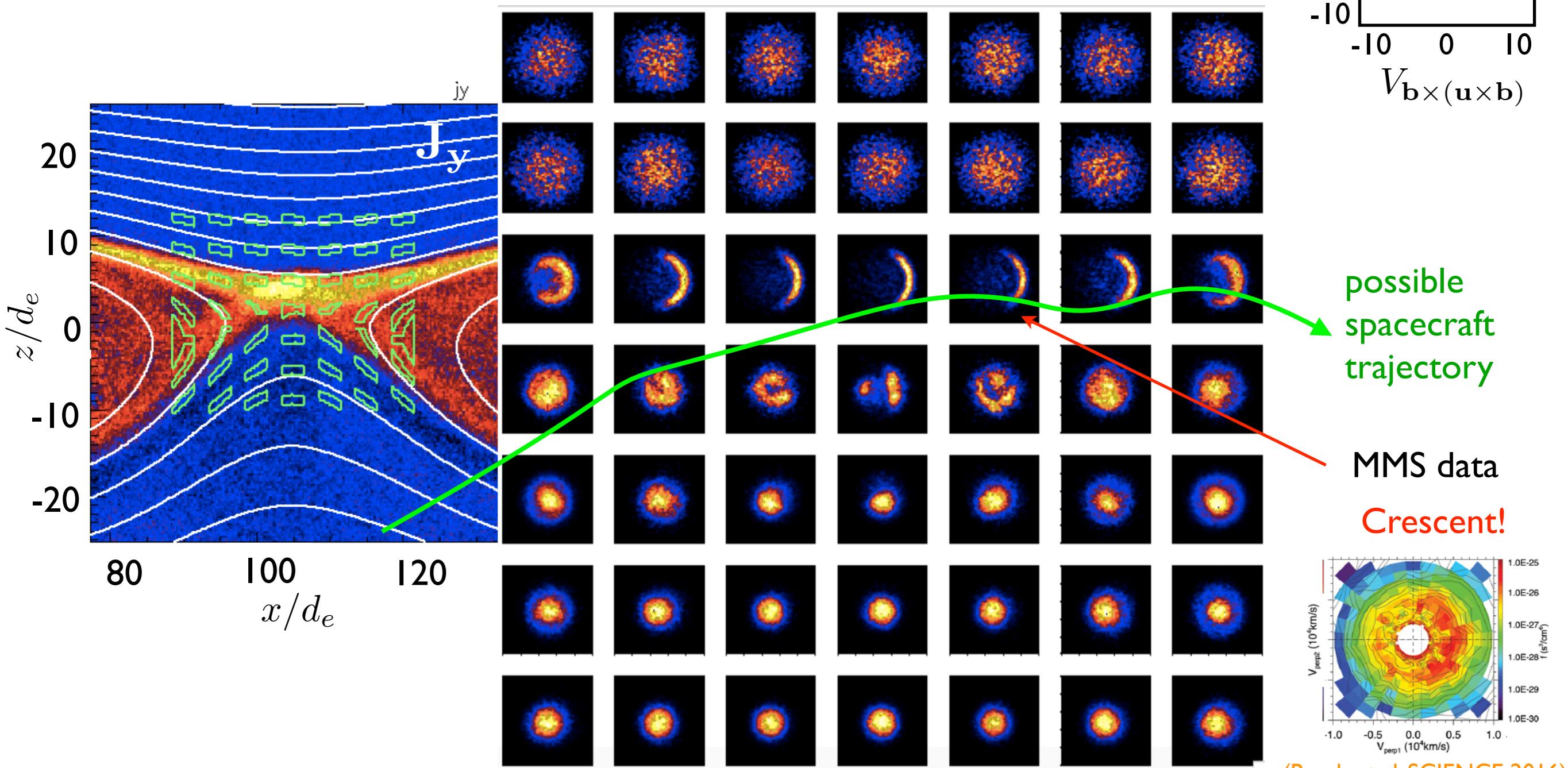
User specification:

- You can request your own boxes.

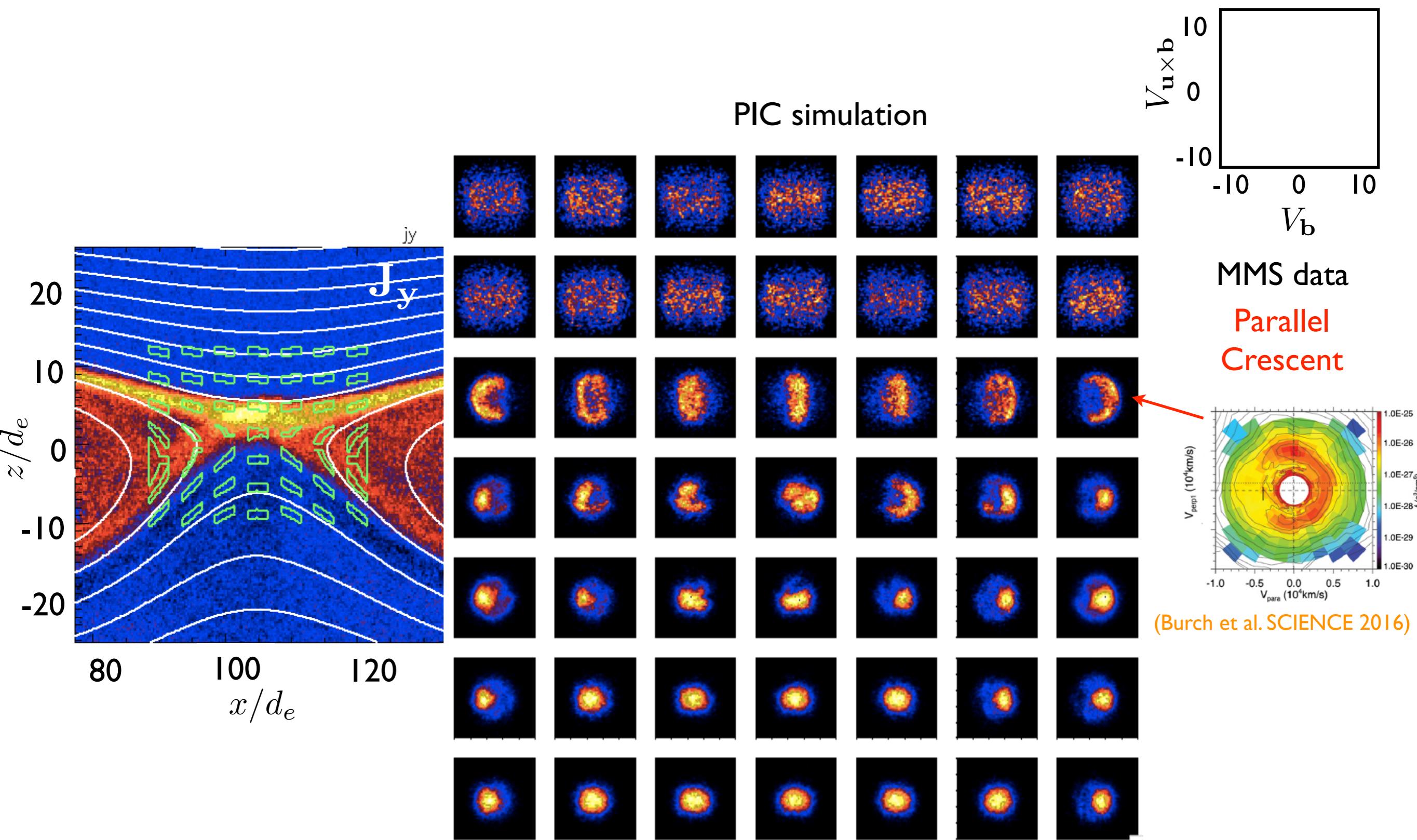
Step 2: We make a run, generate fields, moments & distributions

(e.g., Shuster et al. 2015)

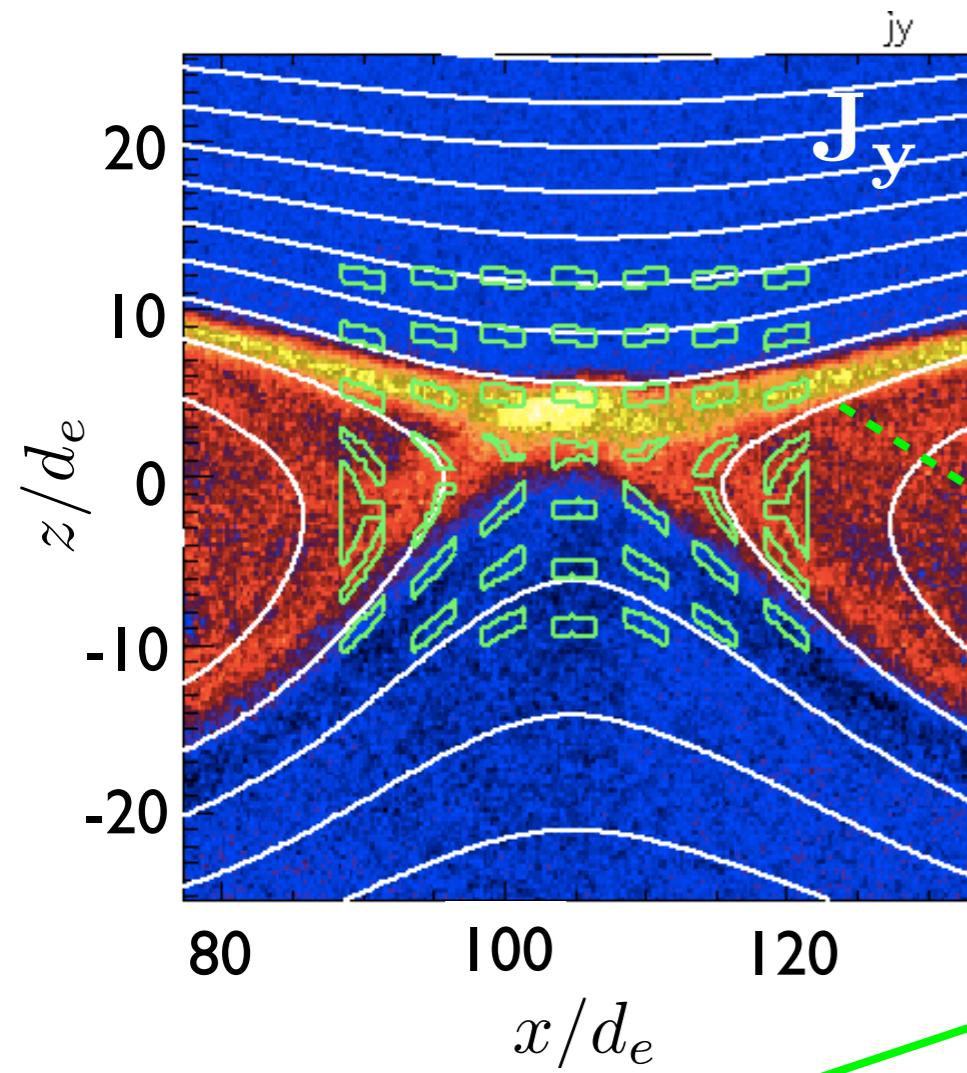
Quick-look



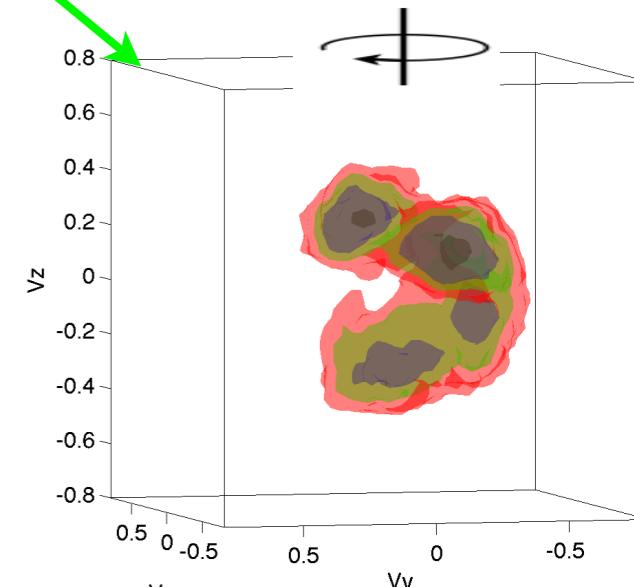
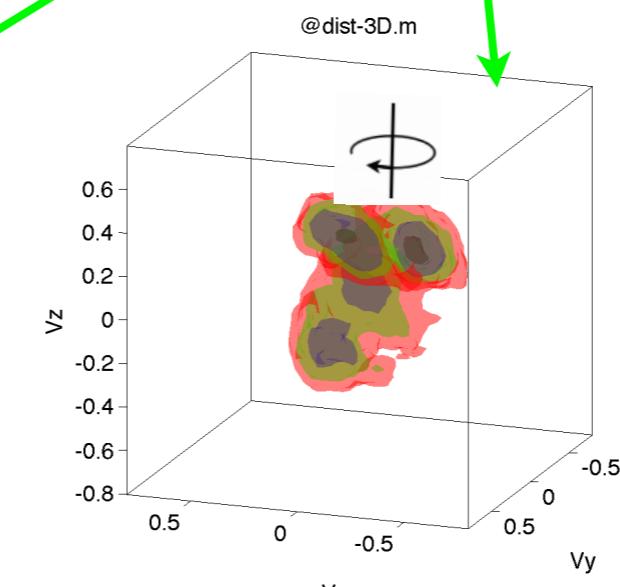
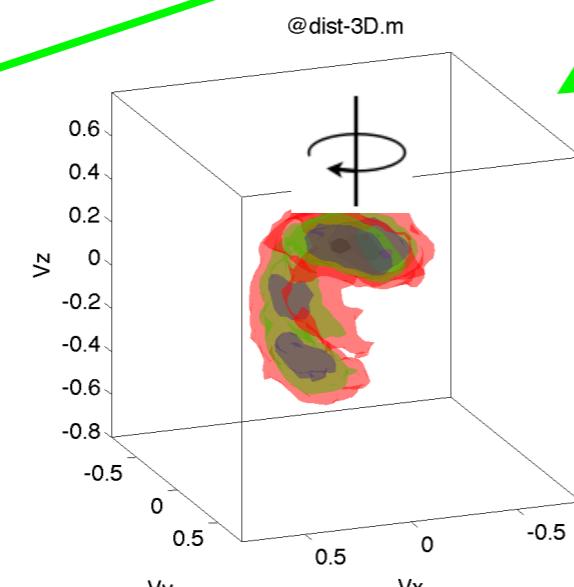
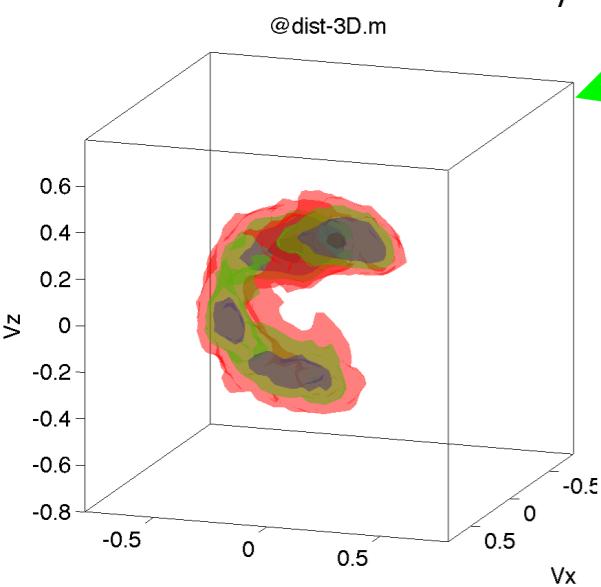
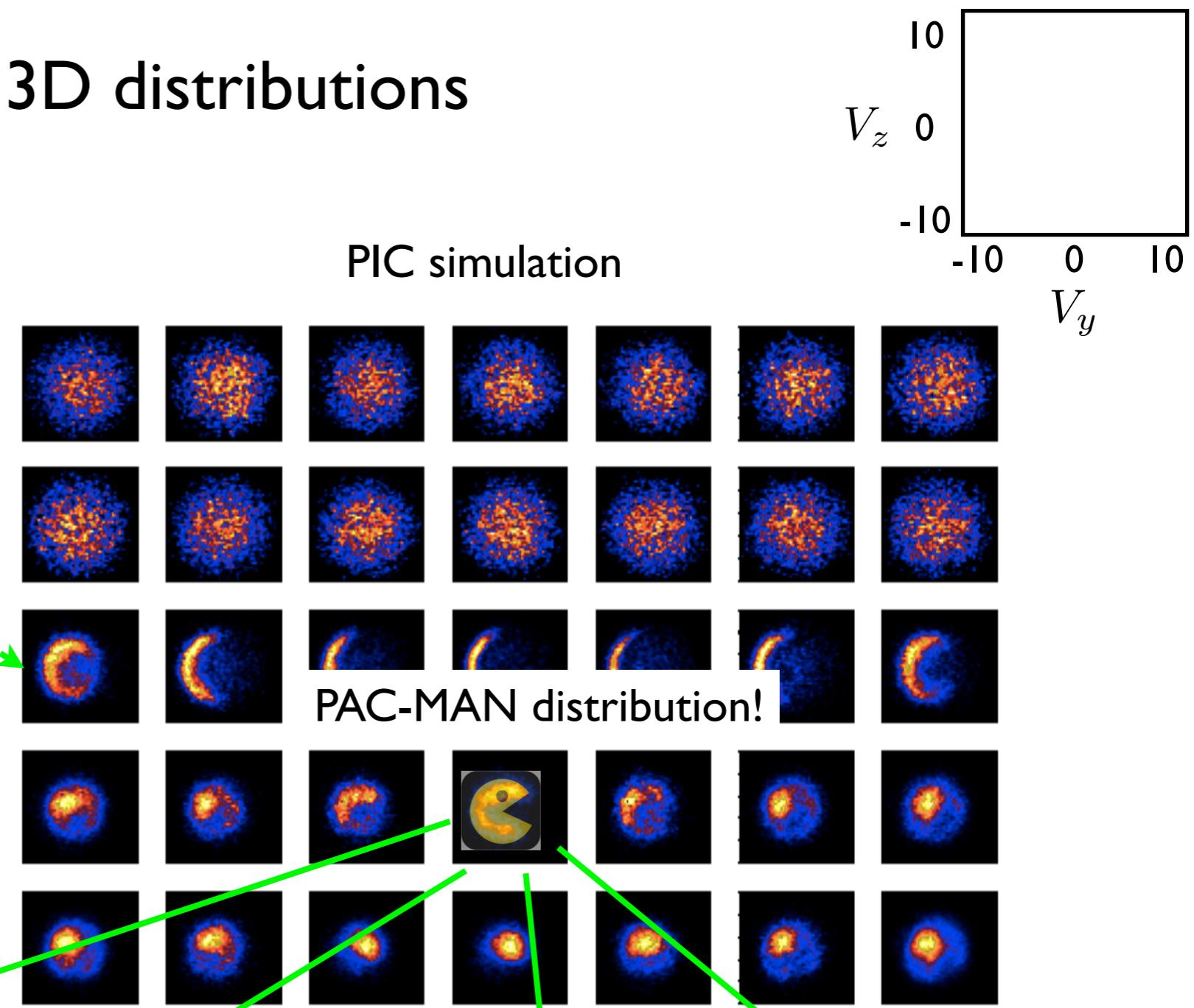
Step 2: We make a run, generate fields, moments & distributions



Full 3D distributions



PIC simulation

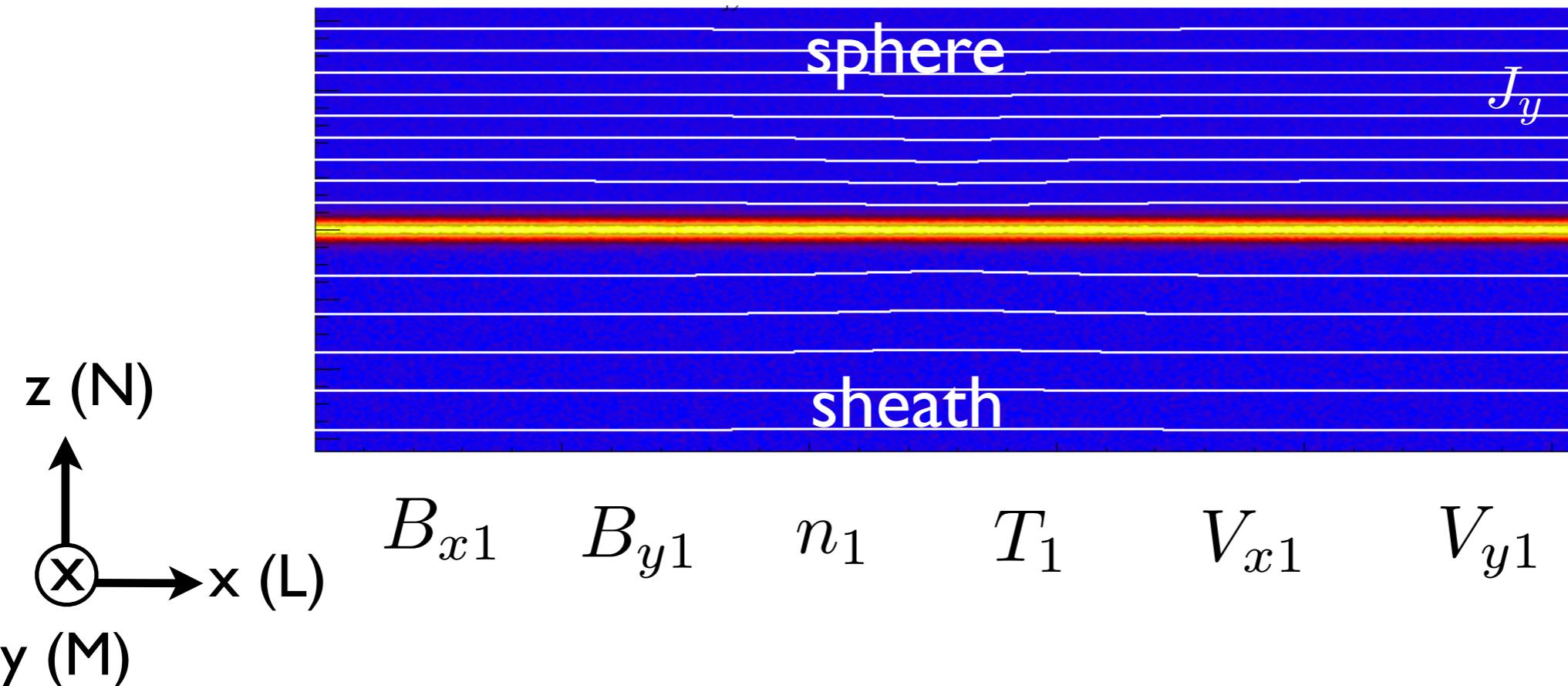


**Run on Request (RoR) of PIC simulations
at the CCMC**

Run on Request

We have generalized the initial condition

$$B_{x2} \quad B_{y2} \quad n_2 \quad T_2 \quad V_{x2} \quad V_{y2}$$



- An user can request a run with customized upstream conditions.
 - CCMC generates particle distribution & field & moment.
 - All data is published on-line & can be analyzed interactively using tools on CCMC.

Profile employed

$$B_x = b_0[1/2 + \alpha_1 \tanh(z/L)]$$

$$B_y = b_0 b_g$$

$$n = \alpha_3[1 - \alpha_2 \tanh(z/L)]$$

$$T = [\alpha_4 - B_x^2/2]/n$$

$$V_x = \alpha_5 \tanh(z/L)$$

$$V_y = \alpha_6 \tanh(z/L)$$

& arbitrary rotation respected to z-axis is possible.

* Note, the pressure balance in the z-direction is always satisfied.

MMS science event (Burch et al., 2016)

Observer-inputs

% in nT
 $B_{1L}=-23;$
 $B_{1M}=-2.278;$
 $B_{2L}=39;$
 $B_{2M}=-2.278;$

% in km/sec
 $V_{1L}=-80;$
 $V_{1M}=-150;$
 $V_{2L}=0;$
 $V_{2M}=0;$

% in cm^{-3}
 $n_1=11.3;$
 $n_2=0.7;$

% in eV
 $T_{e1}=28;$
 $T_{e2}=95;$

$T_i/T_e=10.0;$

Modeler-inputs

$m_i/m_e=25$
 $T_i/T_e=10.0$
 $w_{pe}/w_{ce}=4.0$
 $B_{x2}/B_{x1}=-1.6957$
 $n_2/n_1=0.061947$
 $T_2/T_1=3.3929$
 $B_{g1}=-0.14238$
 $B_{g2}=-0.14238$
 $V_{x_shear}/V_{A1}=-0.26817 \times 2$
 $V_{y_shear}/V_{A1}=-0.50282 \times 2$
 $V_{A1}/c=0.0719$

converting into
code input →

↓ determining
 α_{1-6}

The prototype run size

$m_i/m_e = 25$

$\text{particle}/\text{cell} = 200$

$L_x \times L_z = 51.2d_i \times 25.6d_i$

$n_x \times n_z = 1024 \times 512$

$w_{pe}/w_{ce} = 4.0$

resource required:

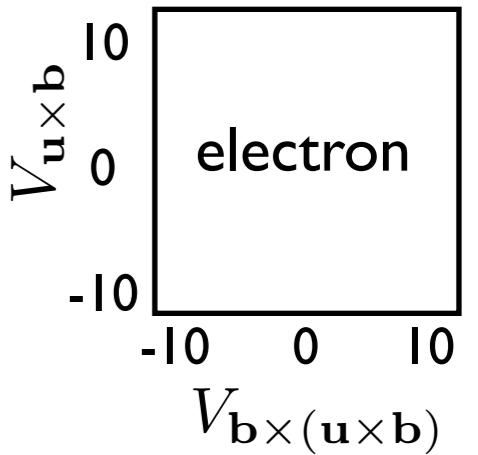
256 CPUs \times 1 hour \sim 256 CPU-hours

Particle data/frame \sim 3 GB

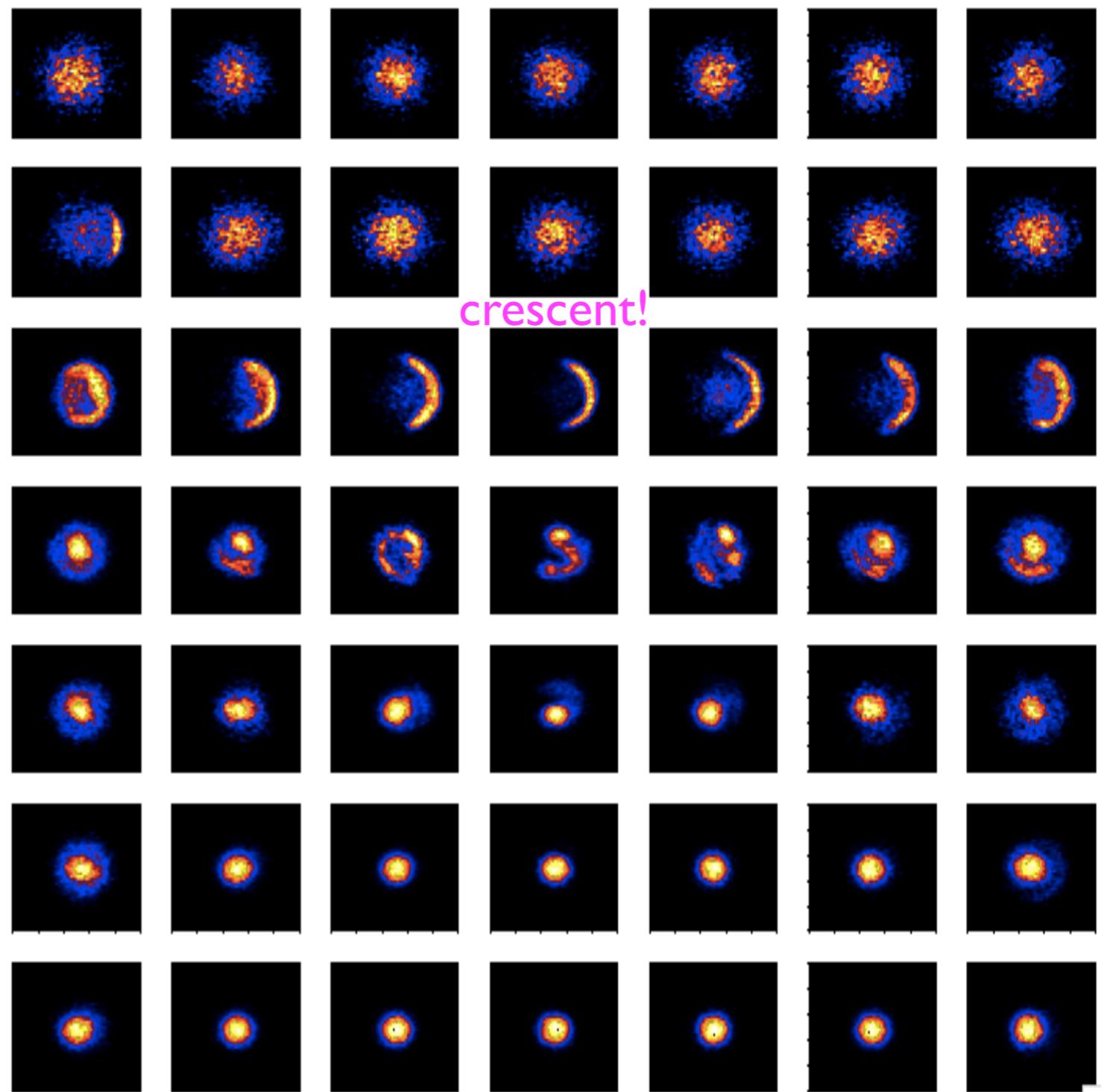
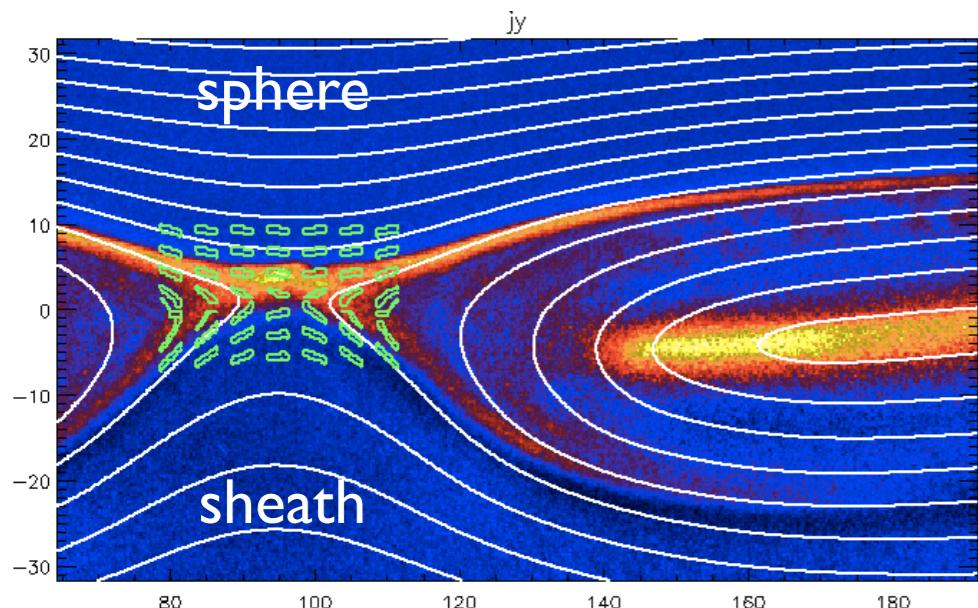
Total fields & moments data \sim 3 GB

A customized run for the event in the SCIENCE paper

[http://ccmc.gsfc.nasa.gov/RoR_WWW/PP/PIC_DIST/2016/
Lutz_Rastaetter_20161121_PP_I/Lutz_Rastaetter_20161121_PP_I.php](http://ccmc.gsfc.nasa.gov/RoR_WWW/PP/PIC_DIST/2016/Lutz_Rastaetter_20161121_PP_I/Lutz_Rastaetter_20161121_PP_I.php)

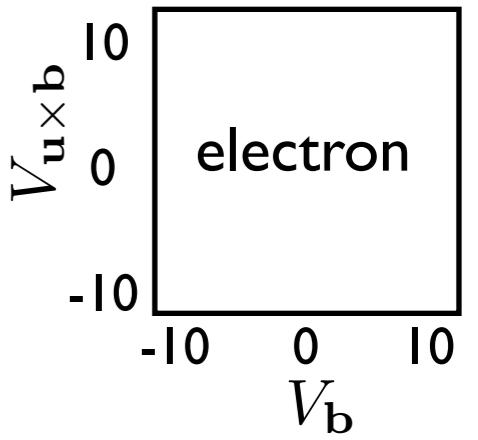


Quick view

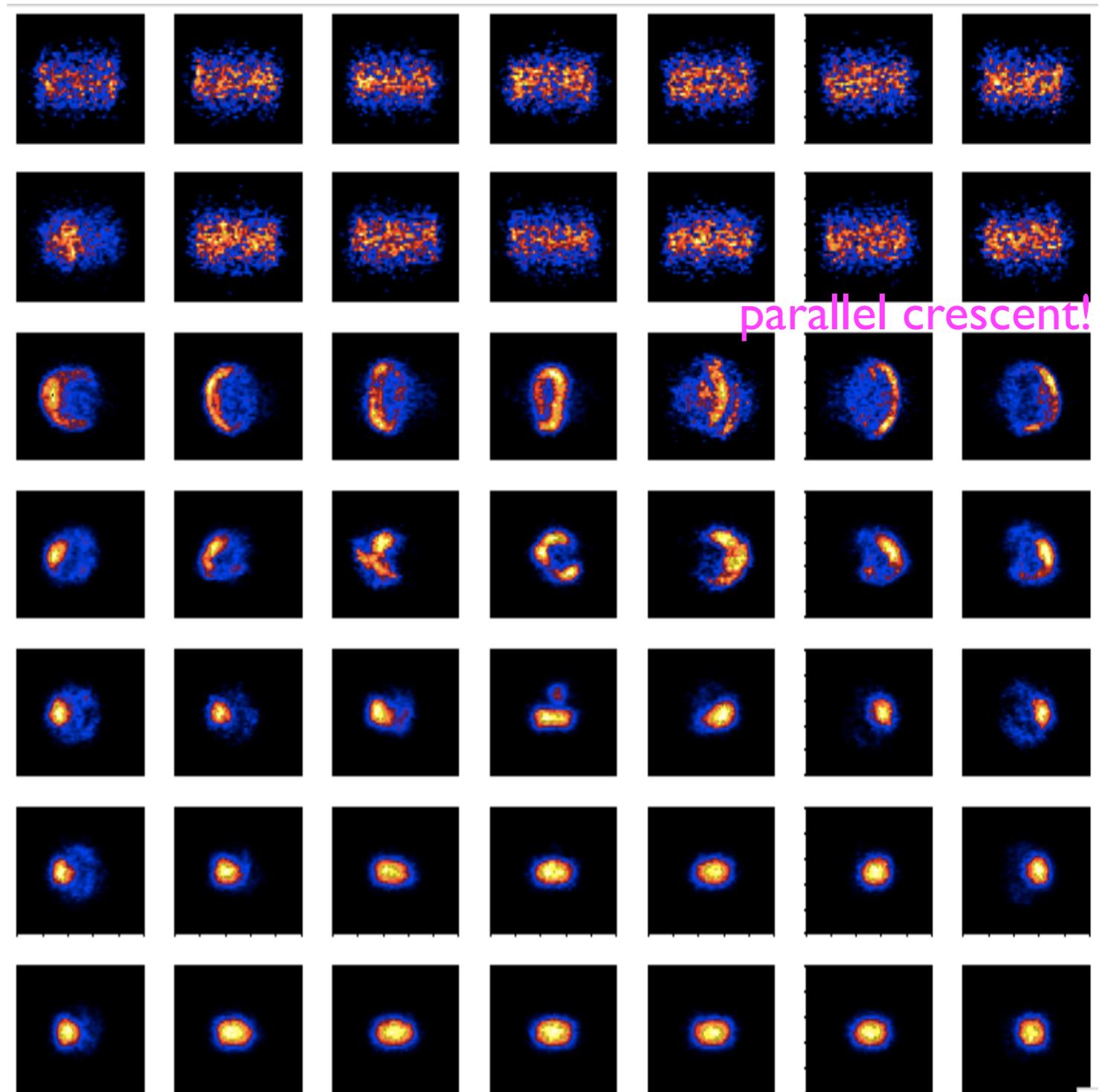
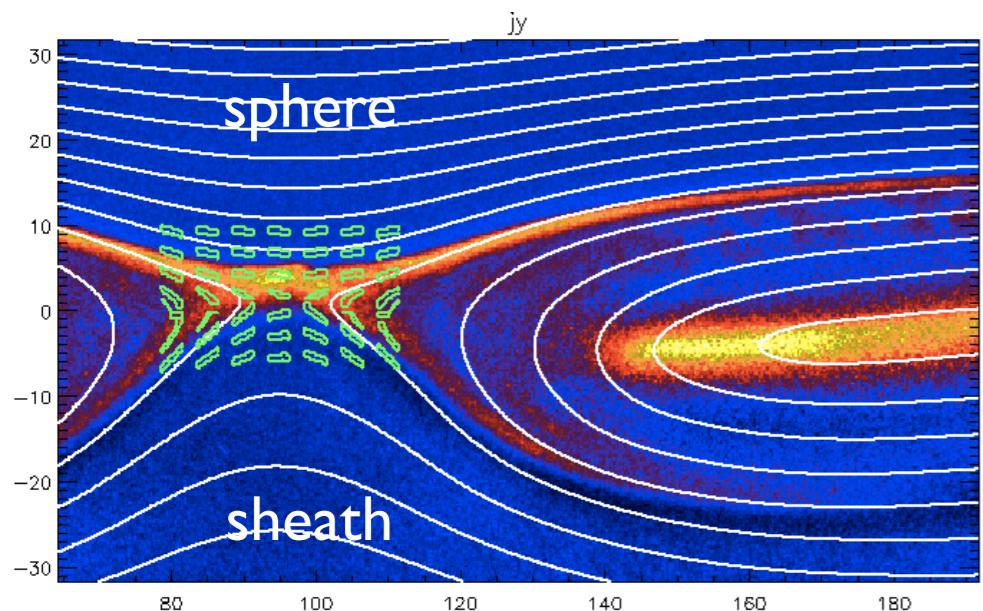


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Quick view



We are accepting order~

- For now, eMail to Lutz.Rastaetter@nasa.gov with the following information

% in nT	% in km/sec	% in $1/\text{cm}^3$	% in eV
B_{1L}	V_{1L}	n_1	T_{e1}
B_{1M}	V_{1M}	n_2	T_{e2}
B_{2L}	V_{2L}		T_i/T_e
B_{2M}	V_{2M}		

we will design & conduct a run, publish it on-line, then notify you.

- Lutz is in the process of designing the on-line interface for the Run on Request (RoR)!

Demo by Lutz